

Flight Programs and X-ray Optics Development at MSFC



M. Gubarev, B. Ramsey ,S.L. O'Dell, R. Elsner, K. Kilaru, C. Atkins, D. Swartz, J. Gaskin and Martin Weisskopf
NASA Marshall Space Flight Ctr., Huntsville, AL 35812, USA

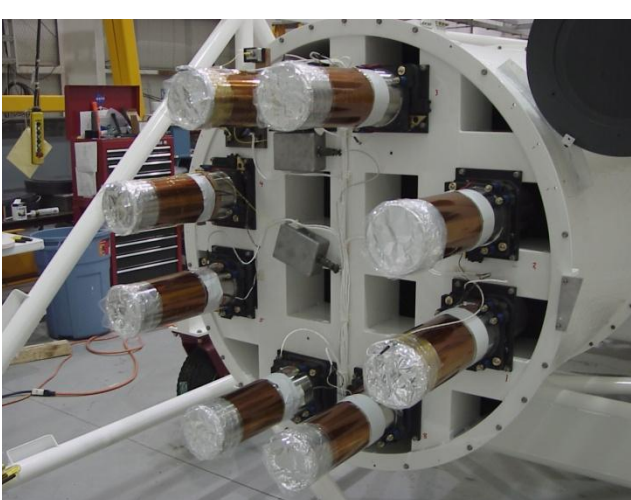
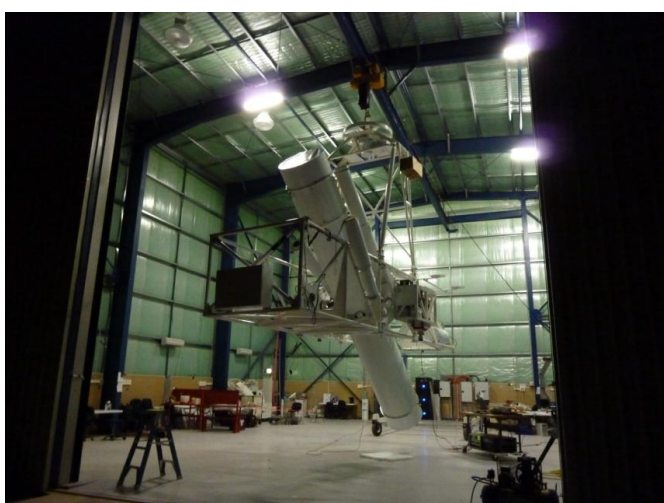
ABSTRACT

The X-ray astronomy group at the Marshall Space Flight Center is developing electroformed nickel/cobalt x-ray optics for suborbital and orbital experiments. Suborbital instruments include the Focusing X-ray Solar Imager (FOXSI) and Micro-X sounding rocket experiments and the HERO balloon payload. Our current orbital program is the fabrication of a series of mirror modules for the Astronomical Roentgen Telescope (ART) to be launched on board the Russian-German Spectrum Roentgen Gamma Mission (SRG.) The details and status of these various programs are presented. A second component of our work is the development of fabrication techniques and optical metrology to improve the angular resolution of thin-shell optics to the arcsecond-level. The status of these x-ray optics technology developments is also presented.

Flight Programs

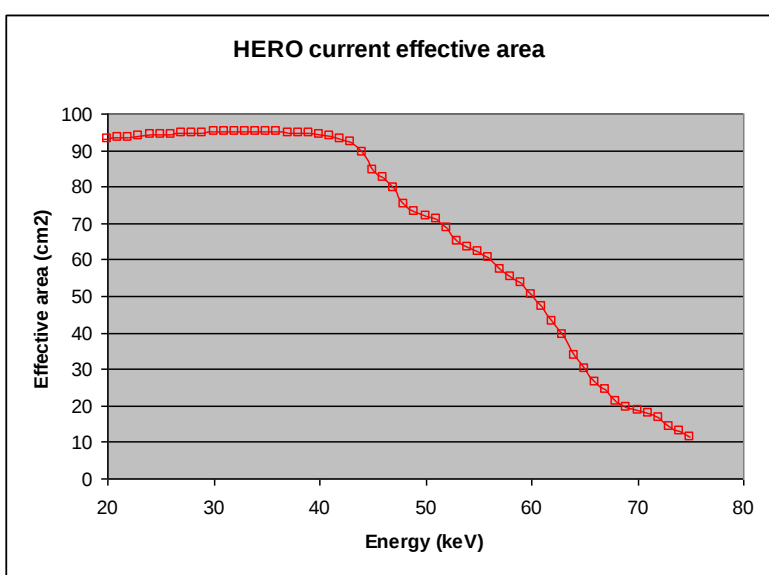


High Energy Replicated Optics (HERO)



HERO, for High Energy Replicated Optics, is a balloon program designed to demonstrate MSFC optics and perform science. Utilizes in-house-fabricated hard-x-ray mirrors plus supporting x-ray detectors, gondola and pointing system.

Number of modules	8
Number of shells	14
Inner, outer diameters	50, 94 mm
Shell thickness	0.25 mm
Focal length	6 m
Angular resolution	~ 13-15 arcsec HPD shells 23 arcsec modules
Field of view	9 arcmin at 40 keV

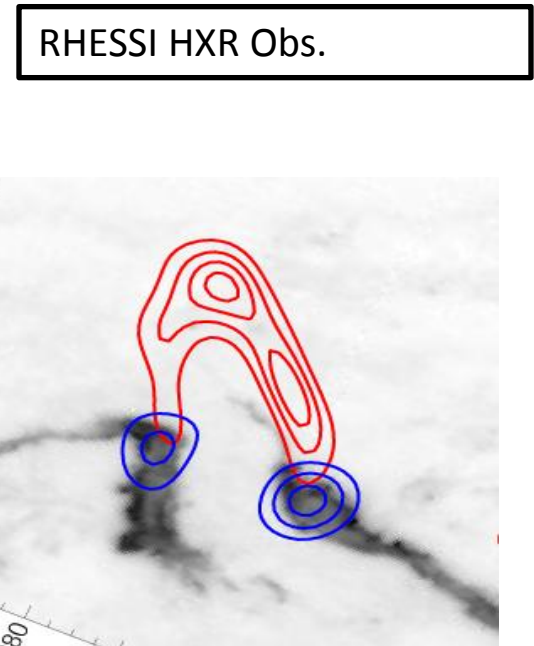


FOXSI : Focusing Optics X-ray Solar Imager

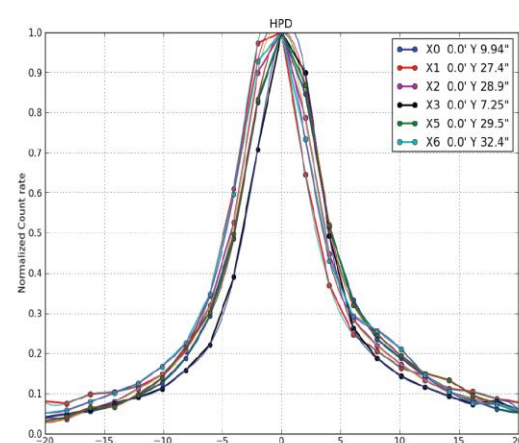
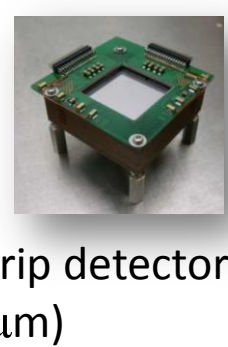
Need to be able to measure weak coronal output against bright footprints, with good angular resolution

Payload Description:

FOXSI is a sounding rocket based payload led by the University of California, Berkeley and consisting of x-ray optics provided by MSFC and focal plane detectors provided by Japan.



Replicated optics telescope module (7 shells & 7 modules)



- Focal Length: 2 m
- Number of modules: 7
- Number of shells: 7 (10)
- Optics Type: Wolter I
- Outer shells radius: 51.51 mm
- Inner shell radius: 37.99 mm (32.48 mm)
- Shell length: 60 cm.
- Energy range: 4 – 15 keV
- Measured angular resolution 22-26 arcsec HPD
- Model fits the measured effective area

Micro-X

Description:

Micro-X is a sounding rocket based payload consisting of x-ray optics (provided by MSFC) and a calorimeter detector led by MIT.

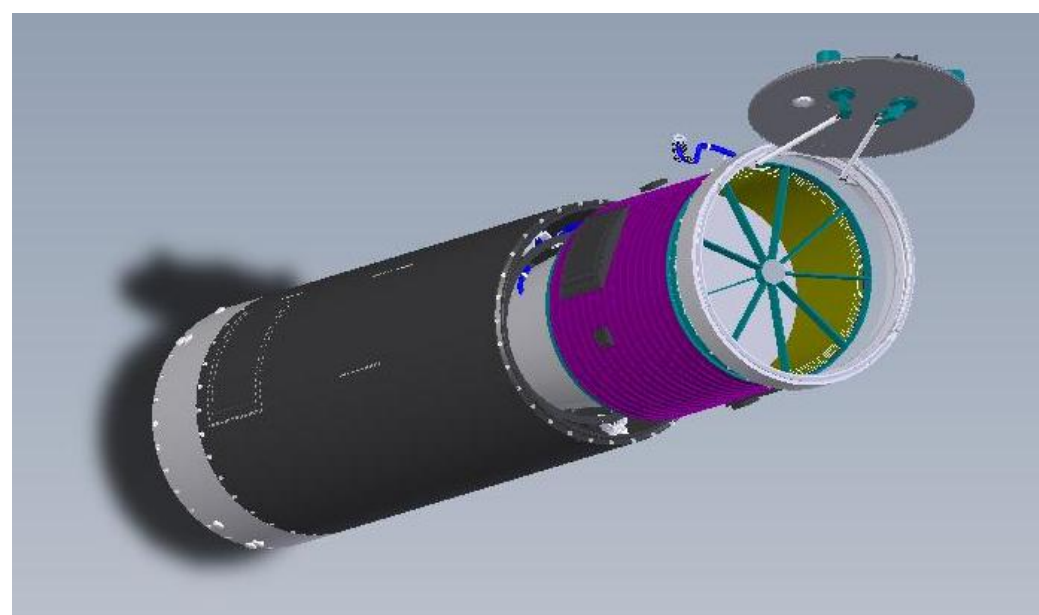
Micro-X will fly in early 2015 and make high-spectral-resolution images of supernova remnants .Puppis A and Cas A.

Design of the 0.5-m-diameter Micro-X optics has just begun.

Customer:

Massachusetts Institute of Technology / Tali Figueroa

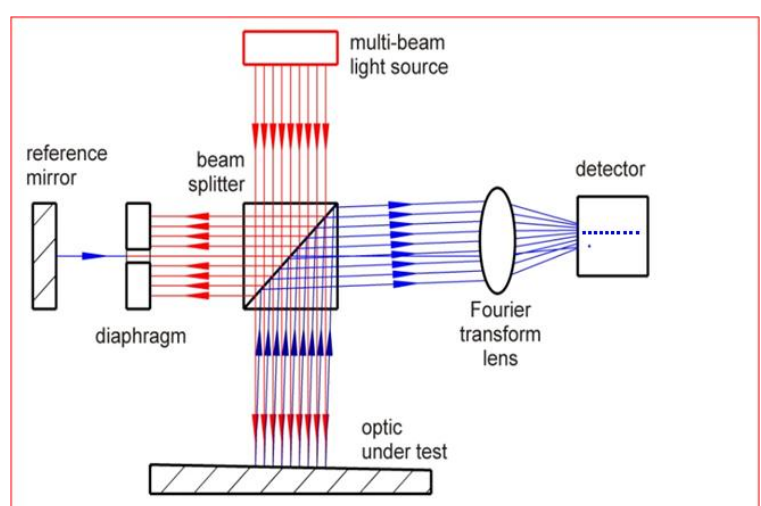
Funded by the Science Mission Directorate, through the low-cost access to space program.



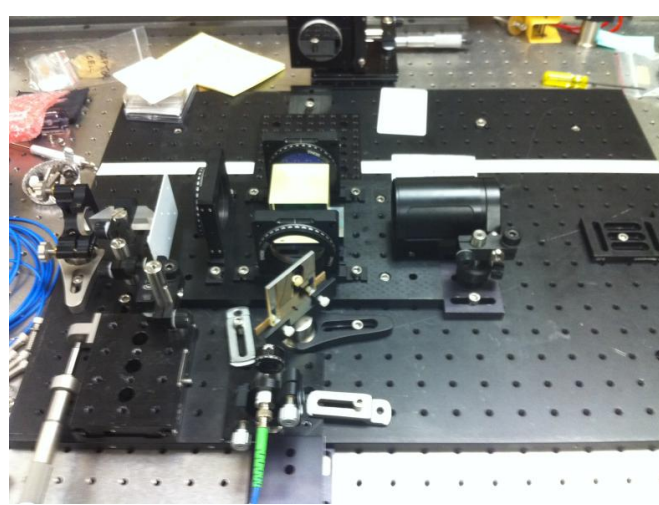
Technology development

Multi-beam long trace-profiler

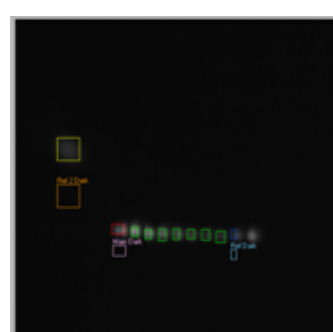
- Make use of advanced technology - Higher resolution and faster 2D detectors, Stable optical sources
- Increase the speed & accuracies of measurements - **Multiple beams**



Schematic of the MBLTP

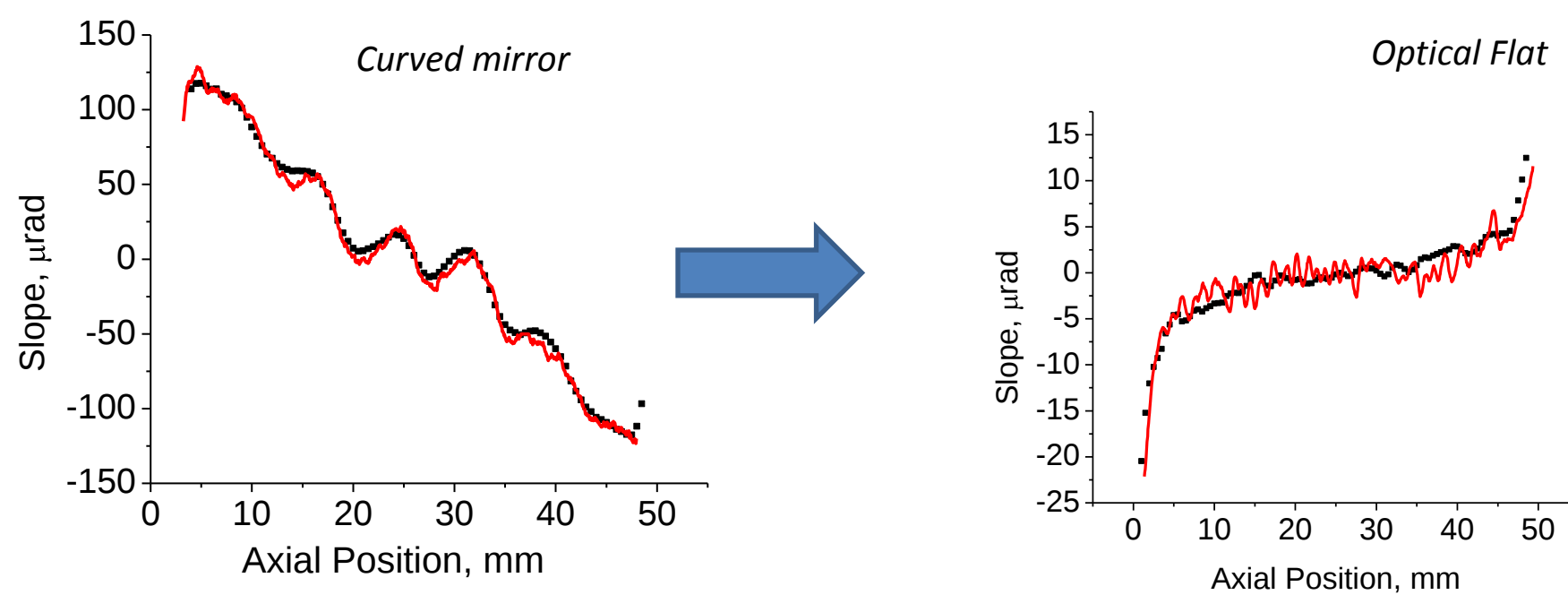


MBLTP breadboard, the detector is not shown (on right)



Screenshot of the detector window. Reference beam is on top left.

- Detector -36 mm x 24 mm area, 7.4 x 7.4 µm pixel size, 1.3 fps for partial frame of 4872x800
- Custom designed FT lens – air-spaced doublet lens, 500 mm focal length, 50 mm diameter.
- The system resolution due to the detector-lens pair is estimated to be ~ 0.23 microrad. Berkeley National Labs has provided software code, we have adapted it for new detector and to track ten beams



- Test have been performed with manual translation stage and thermal enclosure
- Translation by 1 mm
- The slope profiles from the individual beams are stitched

- A multi-beam LTP is under development at MSFC in collaboration with Lawrence Berkeley National laboratory and Brookhaven National Laboratory.
- The components are fabricated, tested individually and assembled into breadboard.
- First tests of the MBLTP has demonstrated the feasibility of the multibeam approach.
- This proof of concept study is believed to form a basis of future modular metrology approach where-in an entire length of the test surface can be measured in a single instance using multiple optical beams.

Astronomical Roentgen Telescope ART

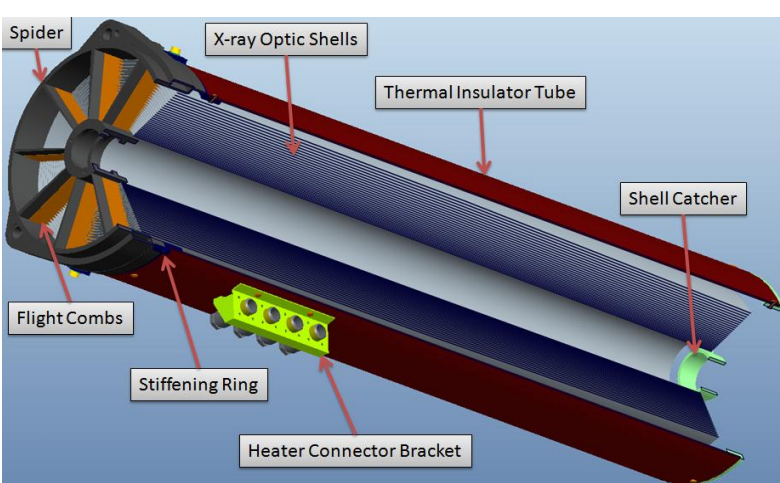
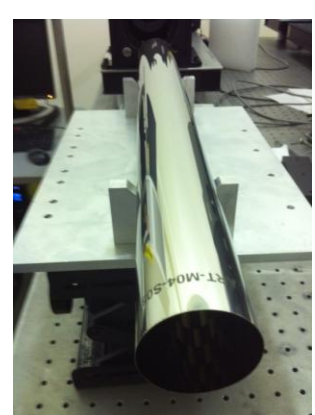


MSFC has designed and is fabricating four ART x-ray optics modules under an International Reimbursable Agreement between NASA and with IKI. The modules have to be delivered to the IKI in June 2013

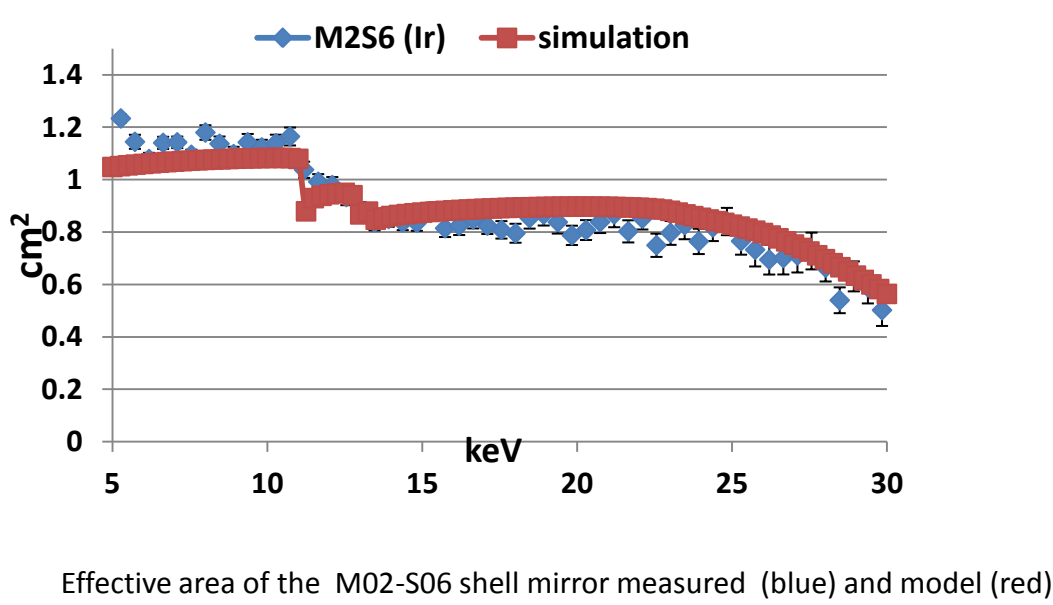
Deliverables: 1 engineering unit (6 shells + mass simulators) and 4 flight units (28 shells/unit)

ART flight module specifications

Parameter	Value
Number of Shells per Module	28
Shell Coating	> 10 nm of iridium (> 90% bulk density)
Shell Total Length, inner and outer diameters	580 mm, 50 mm, 150 mm
Encircled Half Energy Width	Less than 1 mm diameter, center of field of view Less than 2.5 mm diameter, 15 arcmin off axis
Mirror Module Effective Area	≥ 65 cm² at 8 keV (on axis)
Module Focal Length	2700±1 mm
Allowable Total Mass per Module	17 kg including thermal control system
Minimal resonance frequency	40 Hz
Operating Temperature Range	17° C to 23° C



ART mirror module design



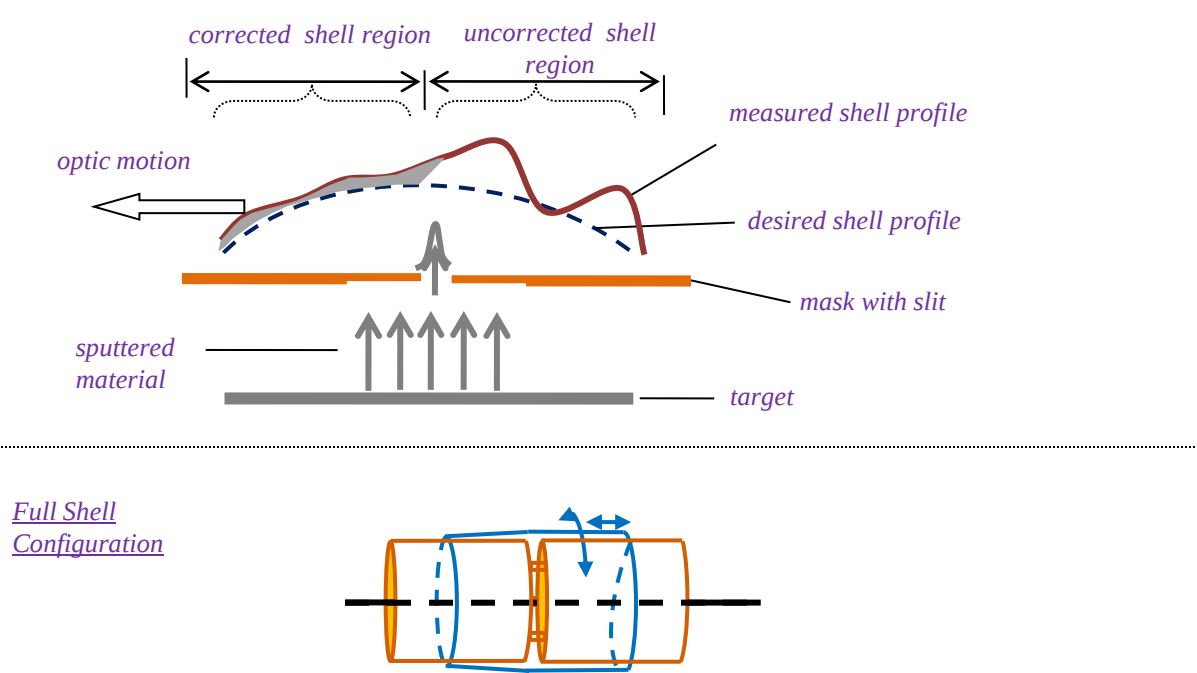
Test was performed at the MSFC's Stray Flight facility

Shell M02-S06 has been tested, the angular resolution is found to be 30 arc seconds HPD, which would correspond to the module performance of ~35 arc seconds

The effective area of the shell exceeds requirements

Improving X-ray optics through Differential Deposition

Addressing profile deviations through differential deposition



Full Shell Configuration

Proof of concept on few-cm-scale medical imaging optics

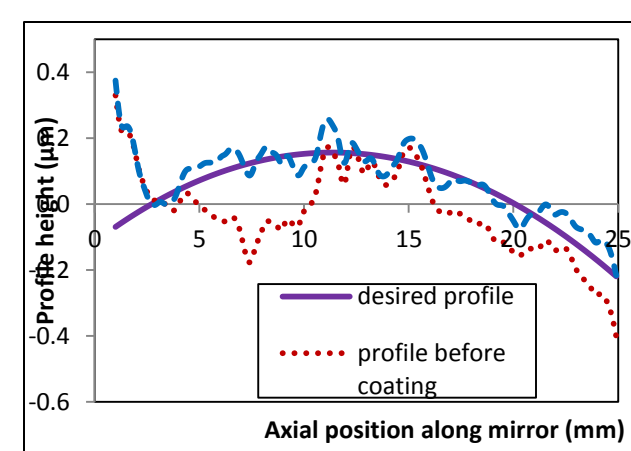
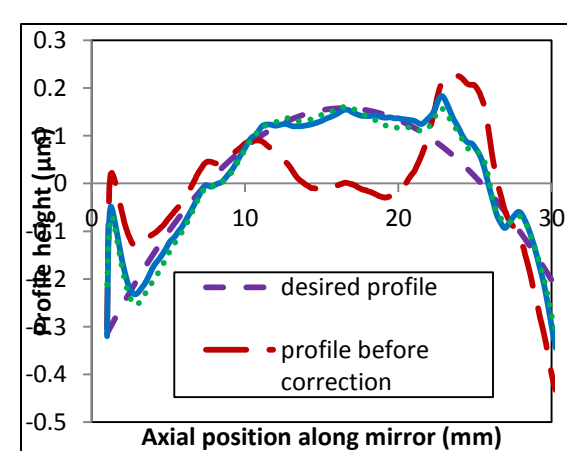
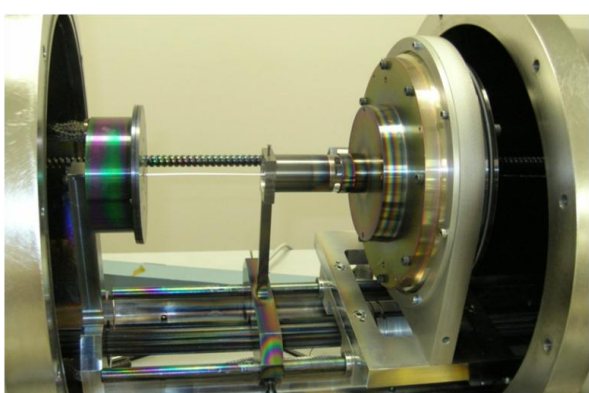


Figure error improvement from 0.11 µm to 0.058 µm rms

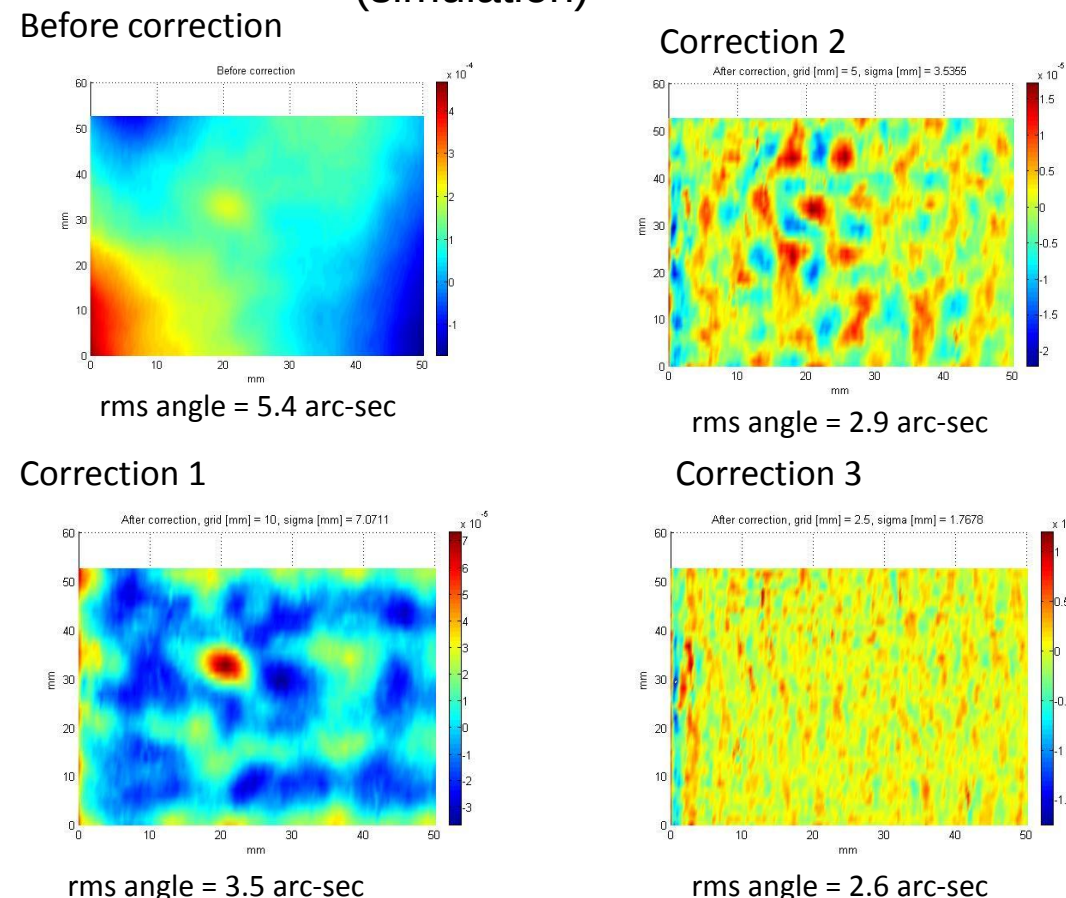


Slope error improvement from 12 arc sec to 7 arc sec rms



Miniature medical optics

Early results for segmented optics (simulation)



Future work will look at addressing:

- Using finer grids and smaller Gaussian widths.
- Solving for the entire surface rather than just a small portion.
- Solving for custom masks to decrease deposition time.